No. 675

CALCULATION OF POTENTIAL FLOW PAST AIRSHIP BODIES IN YAW

By I. Lotz

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The calculation of the potential flow past an airship in yaw is accomplished, according to Von Karman, by dividing the air stream into a flow parallel to the airship axis and one perpendicular to it, after which the potential flow about the solid of rotation exposed in axial direction can then be determined by a method suggested by Rankine. The axis of the body is superposed by sources and sinks and the intensity so defined that the closed stream surface formed by superposition of parallel flow on this source-sink flow is coincident with the surface of the solid. Von Karman (reference 1) pursued an analogous method for defining the potential set up by the lateral flow. He disposes dipolars** on the axis, so that the intensity of this superposition is again defined by the postulate that for the fluid motion, produced by superposition of parallel and dipolar flow, the surface of the solid is stream surface. Both problems yield integral equations. For approximate solution, he substituted a definite number of pieces with constant superposition for the continuous superposition. A numerical calculation of examples according to this method revealed certain difficulties for solids in yaw with not very slender bows, which finally prompted the use of sources and sinks on the surface of the solid instead of dipolars on the axis, and to compute the field of flow by this means. The experiences collected from the calculations and which should prove useful for other similar problems, form the subject of this report.

I. AXIAL SUPERPOSITION

A. Partially Constant Dipolar Superposition

The axis of the airship is assumed to coincide with axis x, counted from the bow. Cylindrical coordinates are introduced so that

$$z = r \cos\varphi, \qquad y = r \sin\varphi \quad \text{(Fig. 1.)}$$

The solid itself, of length l is assumed given by its meridian curve $r(x)$.

The normal velocities induced on the surface by the parallel flow in the direction of the negative axis z, are proportionate to $\cos\varphi$. Since dipols lying on the axis of the solid also set up on the surface of the solid normal velocities proportionate to $\cos\varphi$, the angle φ does not appear in the integral equation defining dipolar motion μ. This equation rests, as already stated, on the premise that the surface is a stream surface, so that on it the normal velocity, composed of the dipolar motion and the component set up by the flow, disappears. Designating the point of the axis carrying the superposition by ξ*, expressing $\frac{dr}{dx} = r'$ and $W =$ velocity of air flow, this equation for the adherence of flow to a parallel of latitude with abscissa x reads as

$$-\int_0^l \frac{\mu(\xi)\, d\xi}{\sqrt{r^2 + (x-\xi)^2}^{\,3}} + 3\, r^2 \int_0^l \frac{\mu(\xi)\, d\xi}{\sqrt{r^2 + (x-\xi)^2}^{\,5}}$$

$$-3\, r\, r' \int_0^l \frac{\mu(\xi)\,(x-\xi)\, d\xi}{\sqrt{r^2 + (x-\xi)^2}^{\,5}} = 4\,\pi\, W \qquad (1)$$

It is known, and Karman also pointed it out in his summary, that exact substitution of the body by a superposition on the axis of rotation is possible only "when the analytical continuation of the potential function free of singularities in the outer space of the solid can be extended to the axis of symmetry without encountering singu-

*The zero point of ξ coincides with that for x.

lar points." Since the question of when or when not this continuance is possible, has never been answered, numerical calculations are apt to have disagreeable surprises in store.*

    With n ~~spaces~~ *intervals* of constant *doublet* dipolar distribution the flow adheres only in n points (control points) of the meridian curve; that means on n circles of the body periphery. (Fig. 2.) The result is an n system of linear equations for the n unknown dipolar intensities μ_ν. With the νth dipol space reaching from ξ_ν to $\xi_{\nu+1}$ (fig. 3), the control point x (reference 1) for flow** adherence is expressed as

$$\sum_1^n \nu\, \mu_\nu \left[\Delta\, g_\nu + r'\, \Delta\, f_\nu \right] = 4\,\pi\, r^2\, W \qquad (2)$$

whereby

$$\left.\begin{array}{l} \Delta\, g_\nu = g(x, \xi_\nu) - g(x, \xi_{\nu+1}), \\[2mm] \Delta\, f_\nu = f(x, \xi_\nu) - f(x, \xi_{\nu+1}), \end{array}\right\} \qquad (3)$$

where

$$\left.\begin{array}{l} g(x, \xi) = \cos\vartheta\,(2 - \cos^2\vartheta), \\[2mm] f(x, \xi) = \sin^3\vartheta \end{array}\right\} \qquad (4)$$

with

$$\tan\vartheta = \frac{r}{x - \xi} \qquad (Fig.\ 3.)$$

    Karman has computed the functions g and f and they appear as dependent on $\dfrac{x - \xi}{r}$. For point $x = 0$, $r = 0$ equation (2) becomes

$$\frac{1}{2}\sum_1^n \nu\, \mu_\nu \left(\frac{1}{\xi_{\nu+1}^2} - \frac{1}{\xi_\nu^2} \right) - m \sum_1^n \nu\, \mu_\nu \left(\frac{1}{\xi_{\nu+1}^3} - \frac{1}{\xi_\nu^3} \right) = 4\,\pi\, W \qquad (5)$$

where m = limit (r r')
        x → 0
        r → 0

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*In the Jablonowski Prize Essay, 1914, G. Herglotz treated a similar problem: Analytical continuance of potential into the inside of the adhering masses, although the fundamental difference of the problems makes it impossible to expect any explanation.
**Manifested in (1) by inserting partially constant superposition μ.

m is a finite number, because r ordinarily goes to zero
as $c\sqrt{x}$. Then the product* (r r') for small x is

$$r\ r' = \frac{c^2}{2};$$

that is, finite for $x \to 0$ and equal to the radius of the
curved sphere for point $x = 0$, $r = 0$.

A different method of visualizing the dipolar inten-
sities is to cut the solid parallel to its axis and perpen-
dicular to the parallel flow and stipulate that the flow
through the outside of this section, generated by the di-
polar motion be equal to the flow of the undisturbed stream
through this area, i.e., equivalent to the section multi-
plied by the flow velocity at infinity. It is not advisa-
ble to effect n such cuts to define the n dipolar in-
tensities, because the slender forms of interest to us
render the exact graphic analysis of the section difficult
and at the same time becloud the dipolar effect when the
cut is made at a distance r contiguous to r_{max} from
the axis. But it is advantageous to occasionally replace
one or two equations with such flow conditions in the sys-
tem (2).

The length of the constant dipolar intensities must
be decided at the very beginning. Knowing μ to be about
proportional to r^2, according to Von Karman, the length
of the spaces depends on the intensity of the change of r.
It will be less toward bow and stern. The major obstacle
lies in the choice of the starting point of the superposi-
tion, whose distance from the bow may be assumed x_1. For
the sphere the ratio of x_1 to curvature radius m is =
1. Long extended bodies having on the axis a stronger di-
polar motion, the interference velocities outside of the
body increase conformably to the flow condition, which
means a shortening of distance x_1. Consequently, the
starting point must be ahead of the center of curvature
of the bow; for the rest, its selection must be left to
experiment.

If a calculation is made first it affords a check on
whether or not point x_1 has been correctly chosen. If
the surface of the body is really a surface of flow then
even the higher differential quotients r'', etc., of the

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* $(r\ r')=\dfrac{d\left(\frac{r^2}{2}\right)}{dx}$; that is, r r' = the derivative of the
curve $\pi r^2 (x)$ divided by 2π.

meridian section of the closed stream surface, produced
by superposed dipolar and parallel flow, must correspond
with those of the surface of the body. We shall see wheth-
er this is the case for the bow. Differentiate equation
(1) according to x and form the limit for $x = 0$ and
$r = 0$. Putting

$$\theta_{p\nu} = \frac{1}{p} \left(\frac{1}{\xi_{\nu+1}^{p}} - \frac{1}{\xi_{\nu}^{p}} \right) \tag{8}$$

and

$$\underset{\substack{x \to 0 \\ r \to 0}}{\text{limit}} \left[(r'^2 + r\, r'') - 1 \right] = s$$

it manifests

$$\sum_{1}^{n} \mu_{\nu} \left\{ s\theta_{3\nu} + 7\, m\, \theta_{4\nu} - 5\, m^2\, \theta_{5\nu} \right\} = 0 \tag{9}$$

The condition is extremely sharp. It shows that even
a minor error in the selection of x_1 is immediately fol-
lowed by a tangible deviation from zero at the right side.
Even an estimate of the anticipated μ values may, under
certain circumstances afford a check on the fulfillment of
this condition. The $\theta_{p\nu}$ values drop very quickly, so
that for the balance only the first spaces come into con-
sideration.

The speedy abatement of the $\theta_{p\nu}$ values with rising
p induces the following: Assuming that very high derivatives
of the bulkhead surface curve are known, it would be
possible to compute μ_1 directly by repeating differentia-
tion of equation (1) and formation of its limit for $x = 0$,
$r = 0$, after having proceeded far enough so that $\theta_{p\nu}$
abates rapidly, so that μ_1, the first term of the sum, only,
need be considered.

Now the course of the dipolar motion for pointed
stern must be included yet. How close to the tip is the
superposition to extend? It is readily shown that inte-
gral equation (1)

$$-\int_{0}^{l} \frac{\mu(\xi)\, d\xi}{\sqrt{r^2 + (x - \xi)^2}^{\,3}} + 3\, r^2 \int_{0}^{l} \frac{\mu(\xi)\, d\xi}{\sqrt{r^2 + (x - \xi)^2}^{\,5}} -$$

$$- 3\, r\, r' \int_{0}^{l} \frac{\mu(\xi)\,(x - \xi)\, d\xi}{\sqrt{r^2 + (x - \xi)^2}^{\,5}} = 4\, \pi\, W,$$

which the dipol motion must fulfill, affords no positive
solution at the right of the maximum section for $r \cdot r' = 0$,
which begins in $x = l$, $r = 0$ as $\mu(\xi) = c(l - \xi)^\lambda$,
where λ and c are positive figures. It means: the di-
polar motion cannot reach the tip and dipolar motion re-
sults only in an approximated rounded-off stern. In a
practical sense there is, of course, no exact tip, but
merely a point of very high, but still finite curvature.
The question of final point of the superposition then
would, provided a certain roundness is desired, revert to
the considerations made for the bow. But the superposi-
tion toward the stern will be so small that a minor vari-
ation in final point will have no effect on the total re-
sult (on the moment, for instance).

B. Continuous, Partially Linearly Variable

Dipolar Superposition

For not very slender solids the dipolar superposi-
tions defined according to the described method are very
much contingent upon the chosen space division. (See figs.
5-8.) This fluctuation admittedly diminishes as the num-
ber of dipolar spaces increases. But for mathematical
reasons a minimum number is desirable, in order to keep
the order of the resolvable equation system low. These
considerations were the reasons for assuming a continuous,
partially linearly variable superposition in place of a
partially constant dipolar motion. The result was, as
proved by examples to be given elsewhere, quite satisfying.

Proceeding from the general equation (1) for flow ad-
herence, we first develop the equation system. With

$$\mu(\xi) = a_\nu + b_\nu \xi \tag{10}$$

in the νth interval, which extends from ξ_ν to $\xi_{\nu+1}$
(fig. 4), equation (1) becomes

$$-\sum_1^n \nu \int_{\xi_\nu}^{\xi_{\nu+1}} \frac{a_\nu + b_\nu \xi}{\sqrt{r^2 + (x-\xi)^2}^3} \, d\xi + 3r^2 \sum_1^n \nu \int_{\xi_\nu}^{\xi_{\nu+1}} \frac{a_\nu + b_\nu \xi}{\sqrt{r^2 + (x-\xi)^2}^5} \, d\xi$$

$$-3rr' \sum_1^n \nu \int_{\xi_\nu}^{\xi_{\nu+1}} \frac{(a_\nu + b_\nu \xi)(x-\xi)}{\sqrt{r^2 + (x-\xi)^2}^5} \, d\xi = 4\pi W$$

or, when writing

$$a_\nu + b_\nu \, \xi = (a_\nu + b_\nu \, x) - b_\nu \, (x - \xi)$$

$$\sum_1^n \nu \, (a_\nu+b_\nu x) \left[-\int_{\xi_\nu}^{\xi_{\nu+1}} \frac{d\,\xi}{\sqrt{r^2+(x-\xi)^2}^3} +3r^2 \int_{\xi_\nu}^{\xi_{\nu+1}} \frac{d\,\xi}{\sqrt{r^2+(x-\xi)^2}^5} - \right.$$

$$\left. -3rr' \int_{\xi_\nu}^{\xi_{\nu+1}} \frac{(x-\xi)\,d\,\xi}{\sqrt{r^2+(x-\xi)^2}^5} \right] - \sum_1^n \nu \, b_\nu \left[-\int_{\xi_\nu}^{\xi_{\nu+1}} \frac{(x-\xi)\,d\,\xi}{\sqrt{r^2+(x-\xi)^2}^3} + \right.$$

$$\left. +3r^2 \int_{\xi_\nu}^{\xi_{\nu+1}} \frac{(x-\xi)\,d\,\xi}{\sqrt{r^2+(x-\xi)^2}^5} -3rr' \int_{\xi_\nu}^{\xi_{\nu+1}} \frac{(x-\xi)^2\,d\,\xi}{\sqrt{r^2+(x-\xi)^2}^5} \, d\,\xi \right] = 4\pi W \quad (11)$$

The integrals of the first sum on the left side appear by Von Karman on page 3; those of the second sum are readily evaluated. Multiplying the whole equation by r^2, yields

$$\sum_1^n \nu \, (a_\nu+b_\nu x) [\Delta \, g_\nu+r' \, \Delta f_\nu] - \sum_1^n \nu \, b_\nu \, r \, [-(\sin\vartheta_{\nu+1}-\sin\vartheta_\nu) +$$

$$+ (\sin^3\vartheta_{\nu+1}-\sin^3\vartheta_\nu) +r'(\cos^3\vartheta_{\nu+1}-\cos^3\vartheta_\nu)] = 4\pi r^2 W \quad (12)$$

For abbreviation, we write:

$$\Delta \, l_\nu = (\sin^3\vartheta_\nu - \sin^3\vartheta_{\nu+1}) -(\sin\vartheta_\nu - \sin\vartheta_{\nu+1}) \quad (13)$$

$$\Delta \, k_\nu = (\cos^3\vartheta_\nu - \cos^3\vartheta_{\nu+1}),$$

so that (12) becomes

$$\sum_1^n \nu \, (a_\nu+b_\nu x) \, [\Delta g_\nu+r' \, \Delta f_\nu] + \sum_1^n \nu \, b_\nu r \, [\Delta \, l_\nu +r' \, \Delta k_\nu] = 4\pi r^2 W$$
$$(14)$$

Carrying the simplification further,

$$\Delta \, g_\nu + r' \, \Delta f_\nu = I_\nu,$$
$$\left. x(\Delta g_\nu + r' \, \Delta f_\nu) + r(\Delta \, l_\nu + r' \, \Delta k_\nu) = II_\nu; \right\} \quad (15)$$

Equation (14), which defines a_ν and b_ν, assumes the form

$$\sum_1^n \nu \, (a_\nu \, I_\nu + b_\nu \, II_\nu) = 4 \pi r^2 W \tag{16}$$

The superposition begins in ξ_1 with value c. Then

$$\left.\begin{aligned}
a_1 &= c - b_1 \times \xi_1 \\
a_2 &= c + b_1(\xi_2 - \xi_1) - b_2 \, \xi_2, \\
&\cdots\cdots\cdots\cdots\cdots\cdots\cdots\cdots\cdots\cdots\cdots \\
a_\nu &= c + b_1(\xi_2 - \xi_1) + b_2(\xi_3 - \xi_2) + \ldots + b_{\nu-1}(\xi_\nu - \xi_{\nu-1}) - b_\nu \xi_\nu
\end{aligned}\right\} \tag{17}$$

From (16) follows

$$\begin{aligned}
c \sum_1^n \nu \, I_\nu + b_1 & \left\{ (\xi_2 - \xi_1) \sum_2^n \nu \, I_\nu - \xi_1 \, I_1 + II_1 \right\} \\
+ b_2 & \left\{ (\xi_3 - \xi_2) \sum_3^n \nu \, I_\nu - \xi_2 \, I_2 + II_2 \right\} \\
+ & \cdots\cdots\cdots\cdots\cdots\cdots\cdots\cdots\cdots \\
+ b_{n-1} & \left\{ (\xi_n - \xi_{n-1}) I_n - \xi_{n-1} I_{n-1} + II_{n-1} \right\} \\
+ b_n & \left\{ \qquad\qquad - \xi_n \, I_n + II_n \right\} = 4 \pi r^2 W,
\end{aligned} \tag{18}$$

which is the equation system for computing the $(n + 1)$ unknown factors $c_1 \, b_1 \ldots b_n$. For the point $x = 0$, $r = 0$, the relation with the notations of equation (8) is

$$\sum_1^n \nu \left\{ a_\nu(\theta_{2\nu} - 3m\theta_{3\nu}) + b_\nu(\theta_{1\nu} - 3m\theta_{2\nu}) \right\} = 4\pi W \tag{19}$$

Putting

$$\left.\begin{aligned}
\theta_{2\nu} - 3 m \, \theta_{3\nu} &= III_\nu, \\
\theta_{1\nu} - 3 m \, \theta_{2\nu} &= IV_\nu .
\end{aligned}\right\} \tag{20}$$

reveals for point $x = 0$, $r = 0$

$$c \sum_{1}^{n} v \ III_v + b_1 \quad \left\{ (\xi_2 - \xi_1) \sum_{2}^{n} v \ III_v - \xi_1 \quad III_1 \quad + IV_1 \right\}$$

$$+ b_2 \quad \left\{ (\xi_3 - \xi_2) \sum_{3}^{n} v \ III_v - \xi_2 \quad III_2 \quad + IV_2 \right\}$$

$$+ \ \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots \ldots$$

$$+ b_{n-1} \left\{ (\xi_n - \xi_{n-1}) \ III_n - \xi_{n-1} \quad III_{n-1} + IV_{n-1} \right\}$$

$$+ b_n \quad \left\{ \quad\quad\quad\quad -\xi_n \quad III_n \quad + IV_n \right\} = 4 \pi W \quad (21)$$

It should be observed, when carrying out the calculation, that the functions g, f, k, l are the same for all bodies dependent on $\dfrac{x - \xi}{r}$;* or, in other words, construe k and l in addition to g and f as in the Karman method. The development of equations (18) and (21) takes somewhat longer than when assuming partially constant dipolar superposition, but the decidedly better results by less number of spaces compensates for this drawback.

C. Examples for A and B

.To illustrate the preceding expositions there follows a compilation of calculated dipolar superposition for bodies which in axial flow are to be substituted by simple singularities on their axis. The body contour is indicated and the computed dipolar superposition plotted in units of the superposition, which corresponds to an indefinitely long cylinder of diameter (2 R) of maximum cross section in a transverse flow. Designating the plotted values with $\mu^e(\xi)$, the dipol moment $\mu(\xi)$ is given by

$$\mu(\xi) = 2 \pi R^3 W \mu^e(\xi)$$

(W = rate of flow). The abscissas of the utilized control points are shown by X's.

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*The functions g, f, k, l were omitted here because of the limited space for the scale, and the comparative ease with which the functions can be plotted.

1. Figuro 5: a body which is produced by a source
and sink of absolutoly equal yiold in a flow parallol to
its connecting lino. The yiold in rolation to the axial
rate of flow is so proportioned that the maximum diameter
equals 1/5 of the distance of the singularitiés. Bow and
stern have the same shape, so the bow is shown only. It
is readily apparent from Figure 5 that the starting point
of the superposition was chosen too far away from the bow.
As a result of its too remote distance from the bow the
superposition of the first spacing is very large and, in
order to neutralize its effects on the other parts of the
contour, a negative superposition must ensue.

2. Figure 6: body produced as in Figure 5; yield of
source and sink such as to make the maximum diameter equal
the spacing of the singularities.

a) The number of dipolar spacings is insufficient.

b) and c) Partially linearly variable steady dipolar
superposition assumed. In case c) it stipulates compli-
ance with differentiated limiting conditions (compare page
5). Note that the fluctuations of the superposition in
a) are perfectly in accord with rate of superposition
b), c).

3. Figure 7: Fuhrmann (reference 2) shape No. IV.
(The sketch shows the source distribution producing the
body in axial flow.) The minor negative dipolar superpo-
sition adjacent to the bow is most probably due to not
quite exact selection of starting point of superposition.

4. Fuhrmann shape No. 1. This body has a very blunt
nose. The dipolar superposition is computed in three
ways. Because of the slight effect of nose and stern in
very slender bodies the dipol superposition can, so to
speak, be effected for two semi-bodies*. For this body
the dipol superposition, computed according to Von Karman's
method, manifests unusual fluctuations at the bow (a).
For that reason the equations, given by the postulated ad-
herence of flow in the first two control points, were re-
placed by two flow conditions, which abated the fluctua-
tions very materially (b). In Figure 8c the starting
point is much farther away from the nose which, being very

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*The minor mutual effect can be accounted for by iteration.

blunt, causes the same phenomena as in example 2.

As concerns bodies with blunt nose, the iteration of the partial equation system for the nose is impossible.

II. SUPERPOSITION ON THE SURFACE

A. Development of Integral Equation for

Surface Superposition and Its Reduction to a

Unidimensional Integral Equation

The hesitation voiced at the beginning of this report against substituting for the body a dipol superposition on the axis, added to the experience gained by computing the cited examples, induced me to apply a different method for defining the potential Φ of the transverse flow around a solid of rotation. Using the same notations as before, namely, x = axis of body, z = negative direction of parallel flow, which is perpendicular to x, and y = direction perpendicular to x and z, potential Φ is defined as follows: It must satisfy the potential equation $\Delta\Phi = 0$, it must act at infinity as $(-W z)$, and on the surface of the body $\frac{\partial\Phi}{\partial n}$ must be $= 0$ (n = direction of the outside normal). Then resolve Φ into Φ_1 and Φ_2, where $\Phi_1 = -W z$. On the surface $\frac{\partial\Phi_1}{\partial n} = -W \frac{\partial z}{\partial n}$. This fulfills the conditions for Φ. Φ_2 must disappear at infinity and the normal derivative on the surface is $\frac{\partial\Phi_2}{\partial n} = +W \frac{\partial z}{\partial n}$. The problem thus resolves to defining Φ_2. This potential was approximated in the Rankine-Von Karman method by the dipol superposition of the body axis. But, conformably to the theory of mass potentials, Φ_2 can equally be represented as potential of an elementary source-sink superposition f spread over the surface of the body. On the premise that the body everywhere has a clearly defined tangential plane, that is, no cusps, this superposition can be expressed by (reference 3): (Grunsel)

$$2\pi f(x,y,z) + \int_S f(\xi,\eta,\zeta) \frac{\cos\psi}{R^2} d\sigma = W \frac{\partial z}{\partial n} \qquad (22)$$

Here   S,   surface of body.

dσ,   surface differential at point ξ, η, ζ.

ψ,   angle between connecting line of fixed sur-
face point $P(x,y,z)$ and running point
$II(\xi,\eta, \zeta)$ with the normal in point $P(x,y,z)$.
(Compare fig. 9.)

R,   distance of fixed from running point.

    The body being symmetrical in rotation, cylindrical
coordinates are introduced:

fixed point P: $r=\sqrt{z^2+y^2}$, $z=r \cos \varphi_r$,   x,

running point II: $\rho=\sqrt{\zeta^2+\eta^2}$, $\zeta=\rho \cos \varphi_\rho$, ξ,   (23)

so that (22) becomes

$$2\pi f(x,r,\varphi_r)+ \int_S f(\xi,\rho,\varphi_\rho) \frac{\cos \psi}{R^2} d\sigma = \frac{W \cos \varphi_r}{\sqrt{1 +\left(\frac{dr}{dx}\right)^2}} \qquad (24)$$

In addition it is written $\frac{dr}{dx} = r'$, $\frac{d\rho}{d\xi} = \rho'$. Extrapolation
then yields from the geometrical relations shown in Figure
9,

$$\cos\psi d\sigma= + \frac{[r-r'(x-\xi)] -\rho \cos(\varphi_r-\varphi_\rho)}{R} \rho d\varphi_\rho d\xi \frac{\sqrt{1+\rho'^2}}{\sqrt{1+r'^2}} \qquad (25)$$

and, with l = length of body,

$$2\pi f(x,r,\varphi_r)+ \int_0^l \int_0^{2\pi} f(\xi,\rho,\varphi_\rho) \frac{[r-r'(x-\xi)] -\rho \cos(\varphi_r-\varphi_\rho)}{R^3}$$

$$\times \frac{\sqrt{1+\rho'^2}}{\sqrt{1+r'^2}} \rho \, d\varphi_\rho \, d\xi = \frac{W \cos \varphi_r}{\sqrt{1+r'^2}} \qquad (26)$$

which, abbreviated, yields:

$$\left.\begin{aligned}
(x - \xi)^2 + r^2 + \rho^2 &= a^2, \\
+ 2 r \rho &= \kappa^2, \\
\varphi_r - \varphi_\rho &= \chi .
\end{aligned}\right\} \qquad (27)$$

So that

$$R^2 = (x - \xi)^2 + r^2 + \rho^2 - 2 r \rho \cos(\varphi_r - \varphi\rho)$$

$$= \alpha^2 - \kappa^2 \cos(\varphi_r - \varphi\rho) = \alpha^2 - \kappa^2 \cos \chi.$$

It is readily seen that $f(x,r,\varphi_r)$ must be proportional to $W \cos \varphi_r$. Consequently, with

$$f(x,r,\varphi_r) = f^e(x,r) \, W \cos \varphi_r,$$

a few changes result in

$$2\pi f^e(x,r) + \int_0^l \int_0^{2\pi} f^e(\xi,\rho) \left[\left\{ r - r'(x-\xi) \right\} \frac{\cos \chi}{\sqrt{\alpha^2 - \kappa^2 \cos \chi^3}} - \right.$$

$$\left. - \rho \frac{\cos^2 \chi}{\sqrt{\alpha^2 - \kappa^2 \cos \chi^3}} \right] \times \sqrt{\frac{1+\rho'^2}{1+r'^2}} \, \rho \, d\chi \, d\xi \, \frac{1}{\sqrt{1+r'^2}} \qquad (28)$$

In this integral equation the integration can be extended over χ. It is necessary to define

$$J_1 = \int_0^{2\pi} \frac{\cos \chi}{\sqrt{\alpha^2 - \kappa^2 \cos \chi^3}} \, d\chi \qquad (29)$$

and

$$J_2 = \int_0^{2\pi} \frac{\cos^2 \chi}{\sqrt{\alpha^2 - \kappa^2 \cos \chi^3}} \, d\chi \qquad (30)$$

These integrations lead, after a number of transformations, to elliptical integrals. Putting (reference 4)

$$\tan \frac{\chi}{2} = t, \qquad \cos \chi = \frac{1 - t^2}{1 + t^2}, \qquad d\chi = \frac{2 \, d t}{1 + t^2}$$

and abbreviating,

$$\lambda^2 = \frac{\alpha^2 - \kappa^2}{\alpha^2 + \kappa^2},$$

yields

$$J_1 = \frac{4}{\sqrt{\alpha^2+\kappa^2}^3}\left[-\int_0^\infty \frac{d\,t}{\sqrt{1+t^2}\sqrt{\lambda^2+t^2}} + \frac{2\,\alpha^2}{\alpha^2+\kappa^2}\int_0^\infty \frac{d\,t}{\sqrt{1+t^2}\sqrt{\lambda^2+t^2}^3}\right],$$

$$J_2 = \left(\frac{\alpha^4}{\kappa^4}\frac{4}{\sqrt{\alpha^2+\kappa^2}^3} - \frac{8\,\alpha^2}{\kappa^4\sqrt{\alpha^2+\kappa^2}}\right)\int_0^\infty \frac{d\,t}{\sqrt{1+t^2}\sqrt{\lambda^2+t^2}}$$

$$+\frac{4\,(1-\lambda^2)}{\sqrt{\alpha^2+x^2}^3}\frac{\alpha^4}{x^4}\int_0^\infty \frac{d\,t}{\sqrt{1+t^2}\sqrt{\lambda^2+t^2}^3} + \frac{4\sqrt{\alpha^2+\kappa^2}}{\kappa^4}\int_0^\infty \frac{d t\sqrt{\lambda^2+t^2}}{\sqrt{1+t^2}^3}$$

Moreover, the introduction of

$$t^2 = \lambda^2\,\frac{\tau^2}{1-\tau^2} \quad \text{and} \quad k^2 = 1 - \lambda^2$$

results, when K denotes the complete elliptic normal integral of the first category, and E that of the second, in

$$J_1 = \frac{4}{\kappa^3\sqrt{\alpha^2+\kappa^2}}\left[\frac{\alpha^2}{\alpha^2-\kappa^2}E(k^2) - K(k^2)\right],$$

$$J_2 = \frac{\alpha^2}{\kappa^2}J_1 + \frac{4}{\kappa^4\sqrt{\alpha^2+\kappa^2}}\left[(\alpha^2+\kappa^2)E(k^2) - \alpha^2 K(k^2)\right].$$

For the sake of simplicity, we put

$$J_1 = \frac{4}{\kappa^2\sqrt{\alpha^2+\kappa^2}}G_1(k^2), \tag{29a}$$

$$J_2 = \frac{4}{\kappa^2\sqrt{\alpha^2+\kappa^2}}\left[G_1(k^2) + G_2(k^2)\right], \tag{30a}$$

that is,

$$G_1(k^2) = \frac{\alpha^2}{\alpha^2 - \kappa^2} \, E(k^2) - K(k^2) = \frac{1}{2} \frac{2-k^2}{1-k^2} \, E(k^2) - K(k^2) \quad (31)$$

$$G_2(k^2) = \frac{2\alpha^2 + \kappa^2}{\kappa^2} \, E(k^2) - \frac{2\alpha^2 - \kappa^2}{\kappa^2} \, K(k^2).$$

$$= 2 \frac{\alpha^2}{\kappa^2} \left\{ E(k^2) - K(k^2) \right\} + \left\{ E(k^2) + K(k^2) \right\}$$

$$= 2 \frac{2-k^2}{k^2} \left\{ E(k^2) - K(k^2) \right\} + \left\{ E(k^2) + K(k^2) \right\} \quad (32)$$

Then (28) becomes

$$2\pi f^e(x,r) + \frac{2}{r} \int_0^l f^e(\xi,\rho) \frac{1}{\sqrt{\alpha^2 + \kappa^2}} \left\{ G_1(k^2) \left[-r'(x-\xi) + (r-\rho) \right] - \rho G_2(k^2) \right\}$$

$$\times \sqrt{\frac{1+\rho'^2}{1+r'^2}} \, d\xi = \frac{1}{\sqrt{1+r'^2}} \quad (33)$$

In this manner the two-dimensional integral equation (22) has been reduced to a unidimensional one, with nucleus or core:

$$\underline{K}(x,r;\xi,\rho) = \frac{2}{r} \frac{1}{\sqrt{\alpha^2 + \kappa^2}} \left\{ G_1(k^2) \left[-r'(x-\xi) + (r-\rho) \right] - \right.$$

$$\left. - \rho \, G_2(k^2) \right\} \sqrt{\frac{1+\rho'^2}{1+r'^2}} \quad (34)$$

This core $\underline{K}$ is examined more closely. It contains the functions $G_1(k^2)$ and $G_2(k^2)$, which are aggregates of ~~whole~~ _{complete} elliptic normal integrals. Both become infinite for $k^2 \rightarrow 1$. Therefore, it is necessary to examine the behavior of the core in the vicinity $k^2 = 1$ more closely. Conformably to the assumptions made previously, a regular fixed point x, r is stipulated (points $x = 0$, $r = 0$ and $x = l$, $r = 0$ are excluded for the time being). The term

$$k^2 = 1 - \lambda^2 = 1 - \frac{\alpha^2 - \kappa^2}{\alpha^2 + \kappa^2}$$

goes, when denoting $(\xi - x)$ by ξ^e, toward 1 as
$1 - \epsilon^2 \xi^{e2}$, where $\epsilon^2 = \left(\dfrac{1 + r'^2}{4\ r^2}\right)$, so long as
$4\ r^2 \gg \xi^{e2}_{max}$.* The term $[- r'(x - \xi) + (r - \rho)]$ goes toward 0 as $-\dfrac{r''}{2}\xi^{e2}\ldots$

Since $K(k^2)$ for $k^2 \to 1$ goes infinite (reference
5) as $-\left[\ln \dfrac{\sqrt{1 - k^2}}{4}\right]$, $E(k^2)$ remains finite, and $[\text{limit}_{k^2 \to 1}$
$E(k^2) = 1]$, it is

and
$$\left.\begin{array}{l}\text{limit}_{k^2 \to 1} G_1(k^2)\ [-r'(x-\xi)+(r-\rho)] = \dfrac{-r''}{4\ \epsilon^2} = \dfrac{-r''r^2}{1+r'^2} \\[4mm] \text{limit}_{k^2 \to 1} \rho\ G_2(k^2) = \text{limit}_{k^2 \to 1} \rho\left[3 + \ln \dfrac{\sqrt{1-k^2}}{4}\right]\end{array}\right\} \quad (35)$$

Thus, core $\underline{K}(x,r;\xi,\rho)$ goes for $k^2 \to 1$ or $\xi^e \to 0$
to infinity as $\ln \dfrac{\epsilon\ |\xi^e|}{4}$, that is, it is integrable.
For the points $x = 0$, $r = 0$ and $x = l$, $r = 0$ $k^2(\xi^e)$
is $= 0$, as far as point $\xi = 0$ or $\xi = l$ itself, where
$k^2 = 1$; that is, $\underline{K}(x,r;\xi,\rho) = 0$ for all points outside
of the fixed point. Consequently, the integral

$$\int_0^l \underline{K}(x,r;\xi,\rho)f^e\ d\xi$$

vanishes for these points. The stern of an airship is usually pointed, obviating the necessity of deliberating whether the limitation imposed previously can be removed; but, as a matter of fact, it cannot. (Reference 6) From the practical point of view this is of no significance, since the stern of an airship has no mathematically exact tip, but merely a point of very pronounced but withal finite curvature.

Thus, abbreviated, as on the preceding page, the integral equation becomes

$$2\pi f^e(x,r) + \int_0^l f^e(\xi,\rho)\underline{K}(x,r;\xi,\rho)d\xi = \dfrac{1}{\sqrt{1+r'^2}} \quad (36)$$

---

*The subscript "max" is to indicate the maximum for ξ^e at which the interpolation formula is still valid.

or $f^e(x,r) = \dfrac{1}{2\pi} \dfrac{1}{\sqrt{1+r'^2}} - \dfrac{1}{2\pi} \int\limits_0^l f^e(\xi,\rho)\,\underline{K}(x,r;\xi,\rho)\,d\xi$ (37)

To resolve it, the well-known iteration method is expedient. Choose as 0th approximation

$$f^e_0(x,r) = \frac{1}{2\,\pi\sqrt{1+r'^2}}\,,$$

define with this f^e_0 the integral

$$f^e\,\underline{K}\,d\xi \approx \int\limits_0^l f^e_0\,\underline{K}\,d\xi$$

and thus compute the first approximation f^e_1. The insertion of this value into the integral yields the second approximation, etc. The convergence of the method is assured (reference 7), because $f^e_0(x,r)$ is finite, the nucleus $\underline{K}(x,r;\xi,\rho)$ is logarithmically infinite at one point only and $\frac{1}{2\pi}$ is less than 1; $f^e(x,r)$ is given by a series of rising powers of $1/\pi$. ~~It becomes~~ When

$$\underline{K}(x,r;\xi,\rho) = \frac{2}{r\sqrt{1+r'^2}}\,\sqrt{1+\rho'^2}\,\underline{K}^e(x,r;\xi,\rho)$$ (38)

and

$$P_n(x,r) = \int\limits_0^l \frac{P_{n-1}(\xi,\rho)}{\rho}\,\underline{K}^e(x,r;\xi,\rho)\,d\xi$$ (39)

is initially put* with

$$P_1(x,r) = \int\limits_0^l \underline{K}^e(x,r;\xi,\rho)\,d\xi$$ (40)

then

$$f^e(x,r) = f^e_0(x,r) - \frac{f^e_0(x,r)}{r}\left[\frac{P_1(x,r)}{\pi} - \frac{P_2(x,r)}{\pi^2} + \frac{P_3(x,r)}{\pi^3} - + \ldots\right]$$

(41)

*For $r = 0$, see preceding page.

B. The Practical Solution of the Integral Equation

a) Numerical definition of core function.- The most time-consuming part of the solution is the definition of the nucleus or core $\underline{K}^e(x,r;\xi,\rho)$ for different fixed points (x,r) of the body. The functions $G_1(k^2)$ and $G_2(k^2)$ being aggregates of whole elliptic normal integrals, are so plotted against k^2, as to be accessible. It should be remembered that both functions $\to \infty$ for $k^2 \to 1$. $G_2(k^2)$ becomes logarithmically infinite, that is, comparatively inferior to $G_1(k^2)$, which approaches infinity as $\frac{1}{1-k^2}$. For more exact reading, it is therefore advisable to plot the function $H(k^2) = (1-k^2)G_1(k^2)$ for k^2 values contiguous to 1, which remains finite for $k^2 \to 1$. Then mark off $H(k^2)$ for a predetermined k^2 and divide this value by $(1-k^2)$, which yields

$$G_1(k^2) = \frac{H(k^2)}{1-k^2}.$$

Figures 10, 11, and 12 show the course of $G_1(k^2)$, $H(k^2)$ and $G_2(k^2)$. The calculus for a predetermined body then begins with the determination of the "form factor" $k^2(x, r;\xi,\rho)$ for a fixed point (x,r) (fig. 13):

$$k^2(x,r;\xi,\rho) = 1 - \frac{\alpha^2 - \kappa^2}{\alpha^2 + \kappa^2} = 1 - \frac{(x-\xi) + (r-\rho)^2}{(x-\xi) + (r+\rho)^2} = 1 - \frac{r_o^2}{r_u^2} \qquad (42)$$

k^2 can be computed by measuring off the distances r_o and r_u. It was found, however, that the purely numerical determination takes about the same time and yields, of course, more accurate data, especially in proximity of (x,r).*

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*k is then written in the form of

$$k^2 = 1 - \frac{1 + \left(\dfrac{r-\rho}{x-\xi}\right)^2}{1 + \left(\dfrac{r+\rho}{x-\xi}\right)^2}$$

Without mathematical term given for the contour, the calculation of

$$Q(x,r;\xi,\rho) = [(r - \rho) - r'(x - \xi)]$$

offers serious obstacles.

This term Q stipulates very exact knowledge of the tangential direction $r' = \dfrac{dr}{dx}$ (fig. 14), and of the second derivation r'', in order to define the limiting value $Q\, G_1(k^2)$ for $k^2 \rightarrow 1$. In Figure 14 Q corresponds to distance AP. It is readily apparent to what errors any inaccuracy in r' of say, 2 per cent, can lead on a long, slender body. The absolute error of Q grows proportionally to $(x - \xi)$. Of advantage is it that k^2 is very small $(G_1(k^2)$ also) for large $(x - \xi)$. In close proximity of x, however, - that is, for small $(x - \xi)$, $G_1(k^2)$ assumes very high values, so that relatively minor discrepancies in Q have a pronounced effect on

$$\underline{K}(x,r;\xi,\rho) = \frac{2}{r}\,\frac{1}{r_u}\left\{G_1\, Q - \rho\, G_2\right\}\sqrt{\frac{1+\rho'^2}{1+r'^2}}$$

because of the existing difference. If the meridian curve of the body is given empirically, a very careful determination of the first and second derivative of r is absolutely necessary. In a given case, it is necessary to again integrate* r'' and r' obtained by numerical differentiation, to ascertain whether the initial values $r(x)$ are attained again - within the permissible limits. If it results in an inadmissible discrepancy $\Delta r(x)$, it is possible in most cases to approximate it by a simple algebraic function and in this manner effect the necessary corrections on r'' and r'.

b) <u>Treatment of point of discontinuity of the core.</u>- After computing the core for a number of fixed points (x,r) the integral $P_n(x,r)$ can be determined. The core being logarithmically infinite for $\xi \rightarrow x$ or $k^2 \rightarrow 1$, integration is possible in interval $\Delta\xi$ also, which contains x, provided f remains finite. Whereas the integration can be made graphically (or planimetrically)

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*The integration of empirical functions is admittedly much safer than the differentiations as calculation method.

outside of this interval, within this critical interval
it must be achieved numerically. (Figure 15.) The plot
shows the curve for $\underline{K}(x,r;\xi,\rho)$. Assume the graphic inte-
gration is made from 0 to $x - \triangle x_l$ and from $x + \triangle x_r$
to l. The choice need not necessarily be $\triangle x_r = \triangle x_l$,
but for expediency, we assume it is, and express it as
$\triangle x$. Then the integral

$$I = \int_{x-\triangle x}^{x+\triangle x} p_n(\xi,\rho)\underline{K}^e(x,r;\xi,\rho)d\xi \qquad (43)$$

must be computed. Herein

$$p_1(\xi,\rho) = 1 \qquad \text{for computing } P_1(x,r)$$
$$p_2(\xi,\rho) = \frac{P_1(\xi,\rho)}{\rho} \qquad " \qquad " \qquad P_2(x,r) \text{ etc.} \qquad (44)$$

For $\underline{K}^e$ its value is written:

$$I = \int_{x-\triangle x}^{x+\triangle x} \frac{p_n(\xi,\rho)}{\sqrt{\alpha^2+\kappa^2}}\left\{G_1(k^2)\left[(r-\rho)-r'(x-\xi)\right]-\rho\,G_2(k^2)\right\}d\xi.$$

For $\xi \longrightarrow x$,

$$G_1(k^2)\left[(r-\rho)-r'(x-\xi)\right]$$

remains finite, as shown elsewhere; thus integral

$$L_1 = \int_{x-\triangle x}^{x+\triangle x} \frac{1}{\sqrt{\alpha^2+\kappa^2}} p_n(\xi,\rho)G_1(k^2)\left[(r-\rho)-r'(x-\xi)\right]d\xi \quad (45)$$

can be graphically determined. There remains

$$L_2 = \int_{x-\triangle x}^{x+\triangle x} \frac{1}{\sqrt{\alpha^2-\kappa^2}} p_n(\xi,\rho)\rho G_2(k^2)d\xi \qquad (46)$$

Since $G_2(k^2)$ approaches infinity as $\ln \frac{\sqrt{1-k^2}}{4}$
when $k^2 \longrightarrow 1$, we put

$$G_2(k^2) = N(k^2) + \frac{4-3\ k^2}{k^2}\ \ln \frac{\sqrt{1-k^2}}{4}$$

The function

$$N(k^2) = E(k^2)\ \frac{4-k^2}{k^2} - \left[K(k^2) + \ln \frac{\sqrt{1-k^2}}{4} \right] \frac{4-3\ k^2}{k^2} \qquad (47)$$

remains finite for $k^2 \rightarrow 1$. (Figure 16.)

Integral L_2 resolves itself to

$$L_2 = L_{21} + L_{22} + L_{23};$$

with

$$L_{21} = \int_{x-\triangle x}^{x+\triangle x} \frac{1}{\sqrt{\alpha^2+\kappa^2}}\ p_n(\xi,\rho)\rho\ N(k^2) d\ \xi \qquad (48)$$

and, taking into account that $k^2 = \frac{4\ r\ \rho}{\alpha^2 + \kappa^2}$ *

$$L_{22} = \frac{1}{2\sqrt{r}} \int_{x-\triangle x}^{x+\triangle x} [p_n(\xi,\rho) - p_n(x,r)]\ \sqrt{\rho} \left(\frac{4-3\ k^2}{k}\ \ln \frac{\sqrt{1-k^2}}{4} \right) d\ \xi$$
$$(49)$$

and

$$L_{23} = \frac{p_n(x,r)}{2\sqrt{r}} \int_{x-\triangle x}^{x+\triangle x} \sqrt{\rho}\ \frac{4-3\ k^2}{k}\ \ln \frac{\sqrt{1-k^2}}{4}\ d\ \xi \qquad (50)$$

L_{21} and L_{22} are to be defined graphically. The integrant of L_{22} remains finite, provided $p_n(\xi,\rho)$ is a continuous function. For expedient evaluation function

$$M(k^2) = \frac{4-3\ k^2}{k}\ \ln \frac{\sqrt{1-k^2}}{4} \qquad (51)$$

is constructed. (Fig. 17.)

Integral L_{23} is numerically interpreted. Developing

$$\left. \begin{array}{l} k^2 = 1 - \epsilon^2\ \xi^{\theta 2} + \gamma\ \xi^{\theta 3} + \ldots, \\ \rho = r + r'\ \xi^{\theta} + r''\ \dfrac{\xi^{\theta 2}}{2} + r'''\ \dfrac{\xi^{\theta 3}}{6} + \ldots \end{array} \right\} \qquad (52)$$

*Follows from easy extrapolation of equation (42).

yields

$$\sqrt{\rho}\ \frac{4 - 3\ k^2}{k} = \sqrt{r}\,(1+\beta_1\ \xi^\theta + \beta_2\ \xi^{\theta^2} + \{\ \beta_3\ \xi^{\theta^3}...) \qquad (53)$$

where

$$\left. \begin{aligned} \beta_1 &= \frac{r'}{2r} \\[2mm] \beta_2 &= 3.5\ \epsilon^2 + \left[\frac{r''}{4r} - \frac{1}{8}\left(\frac{r'}{r}\right)^2 \right] \\[2mm] \beta_3 &= 3.5\ \frac{r'}{2r}\,\epsilon^2 - 3\ \gamma \end{aligned} \right\} \qquad (54)$$

As a rule we stop after the term with ξ^{θ^2} and use the term of the 3d order to estimate the error. Stopping after the terms of the 2d order manifests for L_{23} (50)

$$L_{23} = \frac{p_n(x,r)}{2} \int_{x-\Delta x}^{x+\Delta x} (1+\beta_1\ \xi^\theta + \beta_2\ \xi^{\theta^2})\ln\ \frac{\epsilon\ |\ \xi^\theta|}{4}\ d\xi =$$

$$= \Delta x p_n(x,r) \left[\left(1+\frac{\beta_2}{3}\ \Delta x^2\right)\ln\frac{\epsilon}{4} + (\ln\Delta x - 1) \right.$$

$$\left. + \frac{\beta_2}{3}\ \Delta x^2 \left(\ln\Delta x - \frac{1}{3}\right) \right]$$

The inclusion of the terms of the 3d order in (52) would yield a correction of the order of

$$C_0 = \Delta x p_n(x,r)\ \frac{\gamma}{2\epsilon^2}\left(\beta_1\ \frac{\Delta x^2}{2} + \beta_3\ \frac{\Delta x^4}{5} \right)^{*}.$$

In this manner a reference point is obtained for the selection of Δx.

c) Examples.- The development of the formulas for computing the surface superposition was followed by various illustrative examples in order to check the utility of the method:

1. By a solid of rotation produced at distance $2r_{max}$ by a source and a sink.

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*For computing this correction, we put

$$\ln\ \sqrt{1 - \frac{\gamma}{\epsilon^2}\ \xi^\theta} = -\frac{\gamma}{2\epsilon^2}\ \xi^\theta.$$

2. By Fuhrmann body No. IV.

3. By Fuhrmann body No. 1.

The last was chosen because of its specially inconvenient aspect when computing the axial superposition.

The different approximations of the function of super-position $f^e(x,r)$ have been reproduced in Figures 18-20. The convergence of the method is seen to be quite good. In example 1, the single correction has always the same prefix throughout the entire body length, so that the error caused by the alternating prefixes of the corrective terms (41) is readily estimated. The first correction, of course, produces the principal change in all bodies. Compared to the first calculation of the core, the deter-mination of the higher approximations presents few obsta-cles. For examples 2 and 3, the corrections change the prefix along the body - an effect of the almost conical stern. This change in prefixes is by itself unfavorable for the convergence. However, it being superior for slen-der bodies to that of example 1, the significance of change in prefix is withal subordinate.

C. Calculation of Flow

a) Tangential velocity on the surface of the body.- In practice it is important to be able to compute the pressure on the surface of the body, which means that the velocity must be known at every point of the surface. The velocity is divided into its components - normal and tan-gential. The first, induced by the superposition of the parallel flow in direction of negative axis z over the source flow, is zero. The second is divided in a point (x,r,φ) into components w_1 and w_2 at right angles to one another. Component w_1 clings to the parallel of lat-itude in direction of the tangent, component w_2 in direc-tion of the tangent of the meridian through point (x,r,φ). (Fig. 21.) w_1 is figured positive in direction of grow-ing angle φ; w_2 is valid as positive when pointing from nose to stern. Now, let $f^e(\xi,\rho)$ denote, as before, the surface superposition conjugated to velocity $W = 1$. It results for w_1 and w_2 in

$$w_1 = W \sin \varphi \left[1 + \frac{4}{r\sqrt{1 + r'^2}} \int_0^l f^e(\xi, \rho) \sqrt{1 + \rho'^2} \right.$$

$$\left. \times \frac{\rho}{\sqrt{\alpha^2 + \kappa^2}} G_3(k^2) \, d\xi \right] \tag{55}$$

where

$$G_3(k^2) = \frac{1}{k^2} \left\{ (2 - k^2) K(k^2) - 2 E(k^2) \right\} \tag{56}$$

(fig. 22) and

$$w_2 = W \cos \varphi \left\{ - \frac{r'}{\sqrt{1 + r'^2}} + \frac{2}{r\sqrt{1 + r'^2}} \right.$$

$$\times \int_0^l \frac{f^e(\xi, \rho) \sqrt{1 + \rho'^2}}{\sqrt{\alpha^2 + \kappa^2}} \left[\left((x - \xi) + r'(r - \rho) \right) \right.$$

$$\left. \times G_1(k^2) - \rho \, r' \, G_2(k^2) \right] d\xi \right\} \tag{57}$$

$G_3(k^2)$ is logarithmically infinite, the integral in (55) therefore finite. Note, when evaluating the integral of (57) that, whereas $G_2(k^2)$ is logarithmically infinite, $G_1(k^2)$ is as* $\dfrac{1}{1 - k^2} = \dfrac{1}{\epsilon^2 \, \xi^{e2}}$. On the other hand, since

$$Q^e = (x - \xi) + r'(r - \rho) = - (1 + r'^2) \xi^e - \frac{r''}{2} \xi^{e2} + \ldots$$

changes its prefix for $\xi^e = 0$, $Q^e G_1(k^2)$ goes, when proceeding from the left on x, toward $+\infty$, and toward $-\infty$ when approaching x from the right. The integral

$$\int_0^l \frac{f^e(\xi, \rho) \sqrt{1 + \rho'^2}}{\sqrt{\alpha^2 + \kappa^2}} Q^e G_1(k^2) d\xi$$

must therefore be expressed by its "principal value"

---

* $\xi^e = \xi - x.$

$$\mathop{\mathrm{limit}}_{\delta \to 0} \left[\int_0^{x-\delta} \frac{f^\theta(\xi,\rho)\sqrt{1+\rho'^2}}{\sqrt{\alpha^2 + \kappa^2}} Q^\theta G_1(k^2)\, d\xi \right.$$

$$\left. + \int_{x+\delta}^{l} \frac{f^\theta(\xi,\rho)\sqrt{1+\rho'^2}}{\sqrt{\alpha^2 + \kappa^2}} Q^\theta G_1(k^2)\, d\xi \right]$$

Another point to be noted with w_2 is that the integral for points $x = 0$, $r = 0$ and $x = l$, $r = 0$ does not disappear (in contrast to the normal component set up by source superposition alone). It is best to define the integral for this point separately, foregoing the reductions, which lose their significance because k^2 is always zero up to one point. Tangential velocity $\overline{w}_2$ set up in point $x = 0$, $r = 0$ by source superposition is in direction of negative axis z and known in magnitude by (fig. 23)

$$\left| \overline{w}_2 \right| = 2W \int_{-\pi/2}^{+\pi/2} \int_0^l \frac{f^\theta(\xi,\rho)\cos\varphi_\rho}{R^2} \sin\vartheta \sqrt{1+\rho'^2} d\xi \cos\varphi_\rho\, \rho d\varphi$$

$$= 2W \int_0^l \frac{f^\theta(\xi,\rho)\rho^2\sqrt{1+\rho'^2}}{R^3} d\xi \int_{-\pi/2}^{+\pi/2} \cos^2\varphi_\rho\, d\varphi$$

$$= \pi W \int_0^l \frac{f^\theta(\xi,\rho)\rho^2\sqrt{1+\rho'^2}d\xi}{R^3} \tag{58}$$

wherein $R = \sqrt{\xi^2 + \rho^2}$.

b) The potential in any point of the space.- This is given by

$$\Phi = - 4W \cos\varphi_r \int_0^l \frac{f^\theta(\xi,\rho)\sqrt{1+\rho'^2}}{\sqrt{\alpha^2 + \kappa^2}} G_3(k^2)\rho\, d\xi \tag{59}$$

$G_3(k^2)$ being the function defined by (56).

Note, that $k^2 = 1 - \left(\dfrac{r_0}{r_u}\right)^2$ is always 1, when computing the potential in points outside of the body. As a result, no discontinuity places of the integrand appear.

SUMMARY

An outline of Von Karman's method of computing the potential flow of airships in yaw by means of partially constant dipolar superposition on the axis of the body is followed by several considerations for beginning and end of the superposition. Then this method is improved by postulating a continuous, in part linearly variable dipolar superposition on the axis.

The second main part of the report brings the calculation of the potential flow by means of sources and sinks, arranged on the surface of the airship body. The integral equation which must satisfy this surface superposition is posed, and the kernel reduced to functions developed from whole elliptic normal integrals. The functions are shown diagrammatically. The integration is resolvable by iteration. The consequence of the method is good. The formulas for computing the velocity on the surface and of the potential for any point conclude the report.

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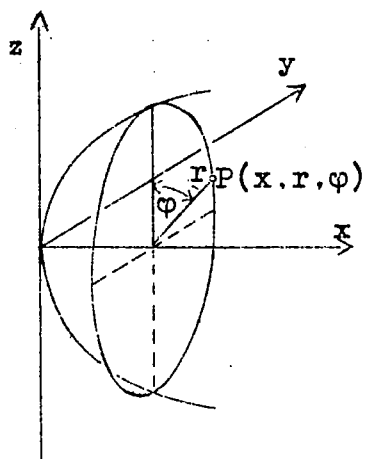
Fig. 1 Coordinate axes
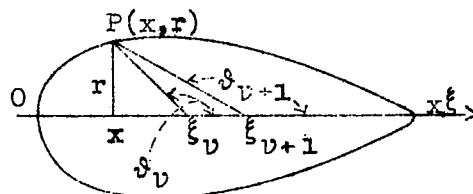
Fig. 3 Notations
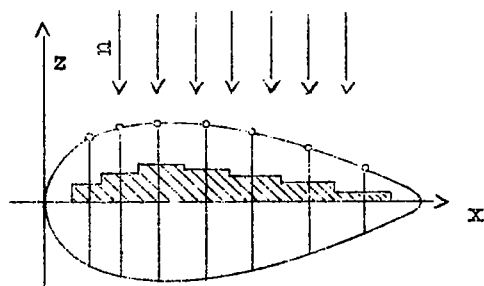
Fig. 2 Body with dipolar super-
        position on the axes.
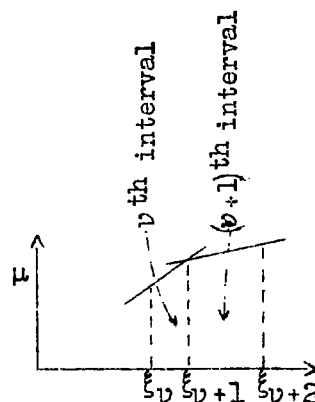
Fig. 4 Intervals by partially
        linearly variable di-
        polar superposition.
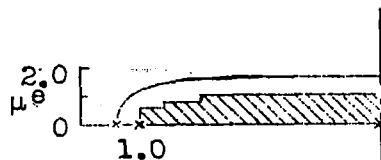
Fig. 5a
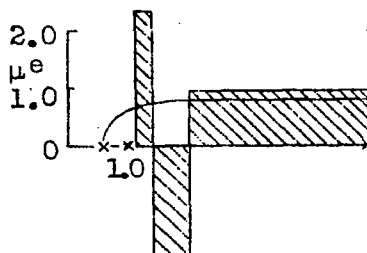
Fig. 5b
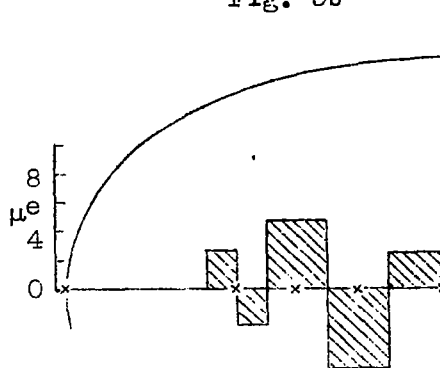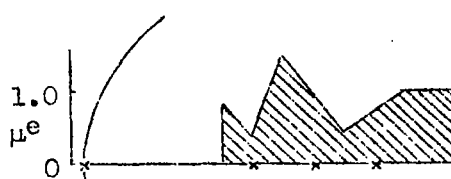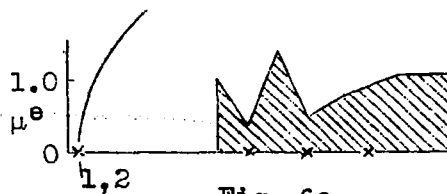
Figs. 5,6
Computed
dipolar
superpositions

Fig. 6a

Fig. 6b

Fig. 6c

Generating source distribution.
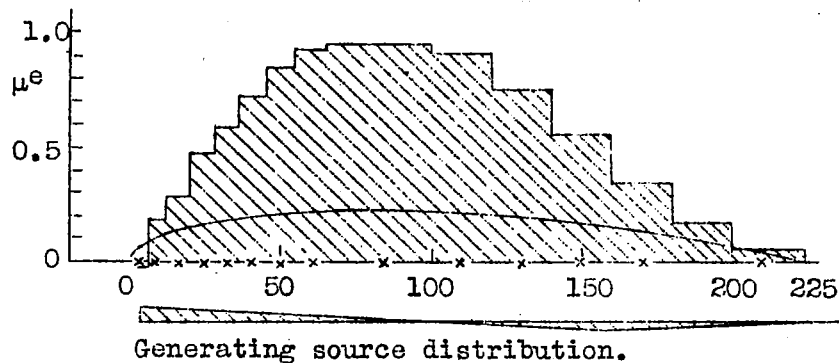
Fig. 7
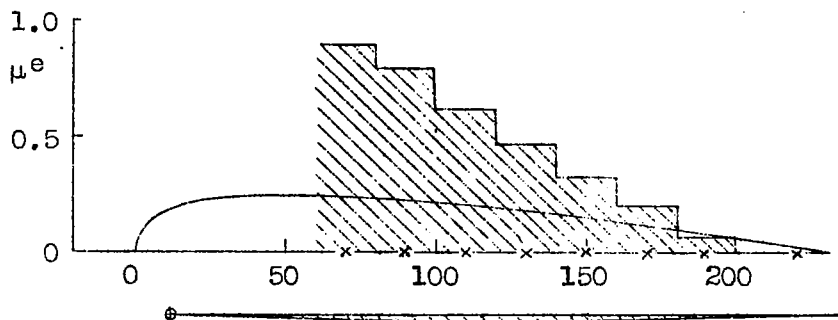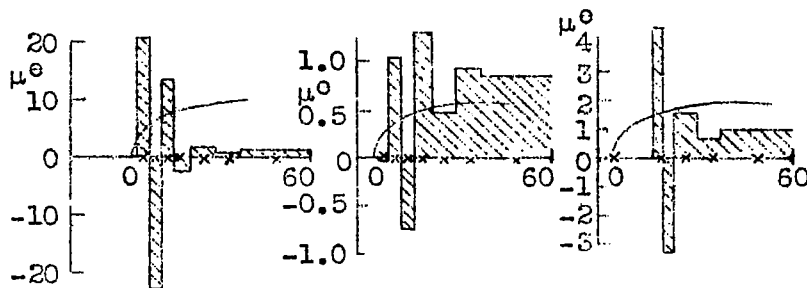
Fig. 8

Fig. 8a   Fig. 8b   Fig. 8c

Figs. 7,8   Computed dipolar superpositions.
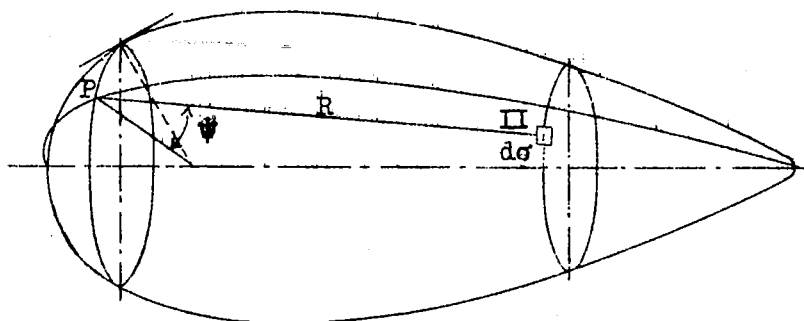
Fig. 9 Notations for surface superposition.
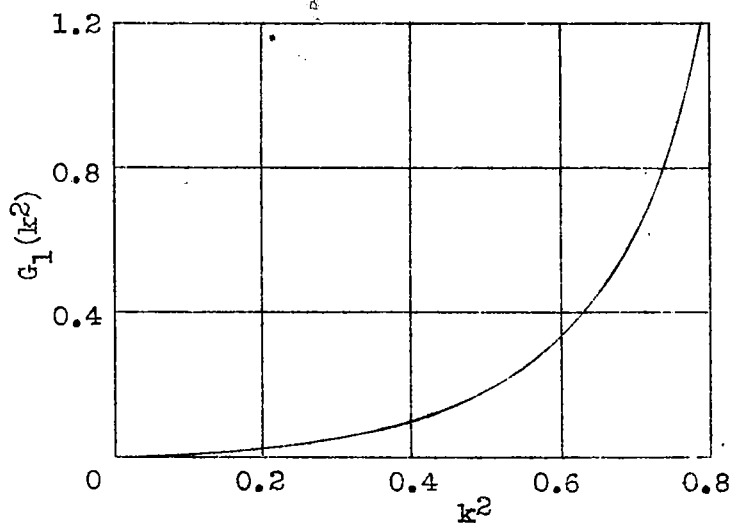
Fig. 10   The $G_1(k^2)$ function.
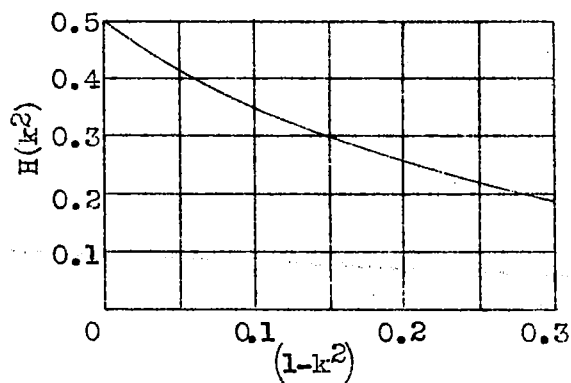
Fig. 11   The $H(k^2) = (1-k^2)G_1(k^2)$ function.
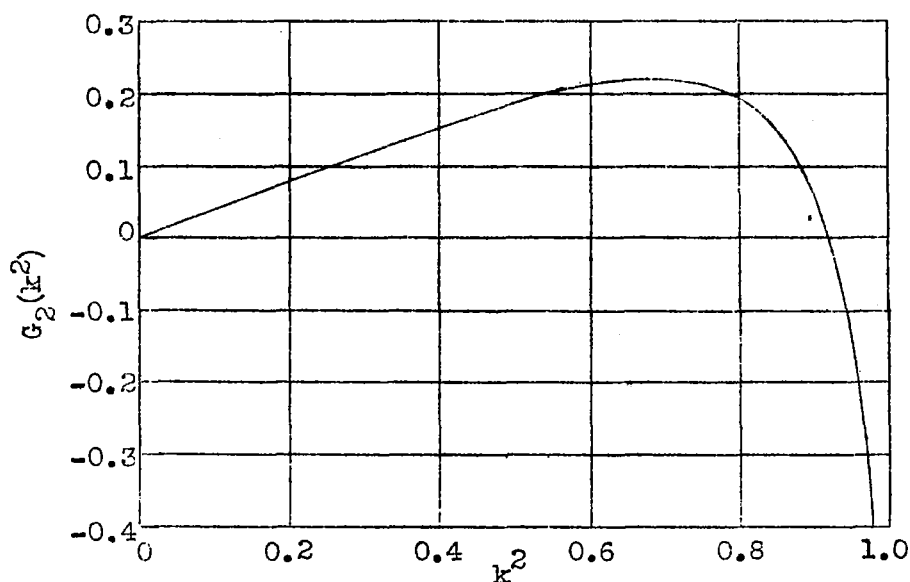
Fig. 12   The $G_2(k^2)$ function.
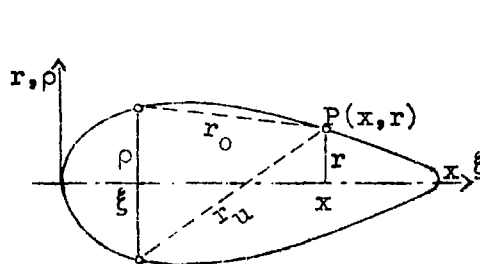
Fig. 13   Definition of form factor k^2
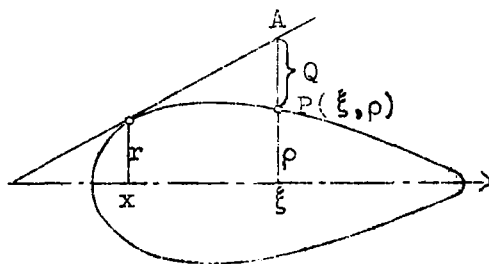
Fig. 14   Geometrical significance of Q
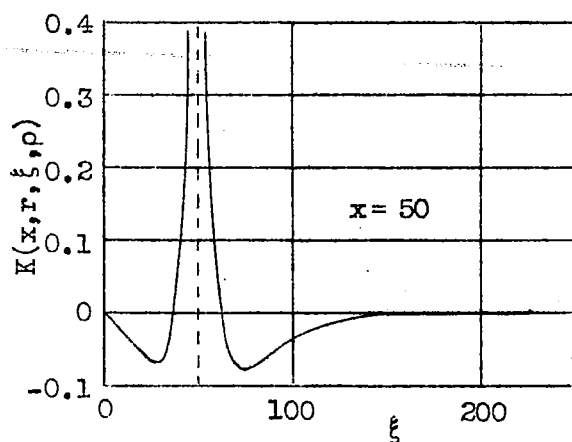
Fig. 15   Function of curve $K(x,r,\xi,\rho)$ for $x = 50$. (Fuhrmann shape No.1, $l = 230$)
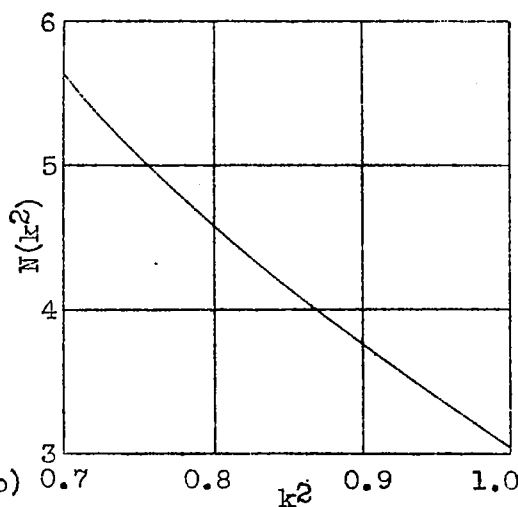
Fig. 16   The $N(k^2)$ function.
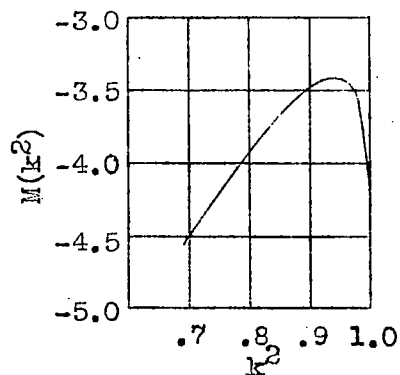
Fig. 17   The $M(k^2)$ function.
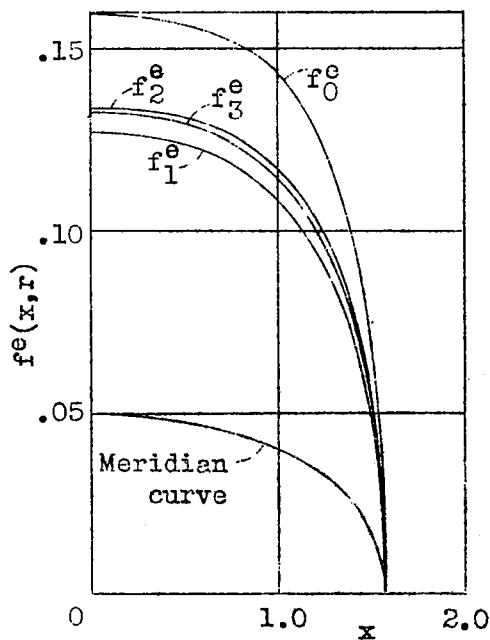
Fig. 18   Surface superposition of source – sink body.
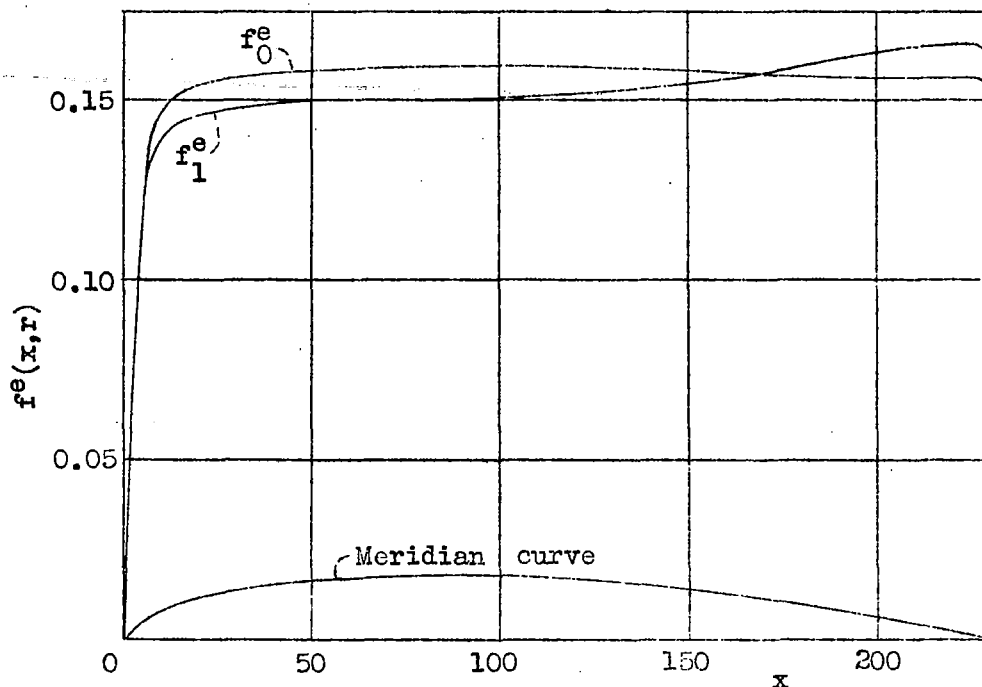
Fig. 19   Surface superposition of Fuhrmann shape No. IV
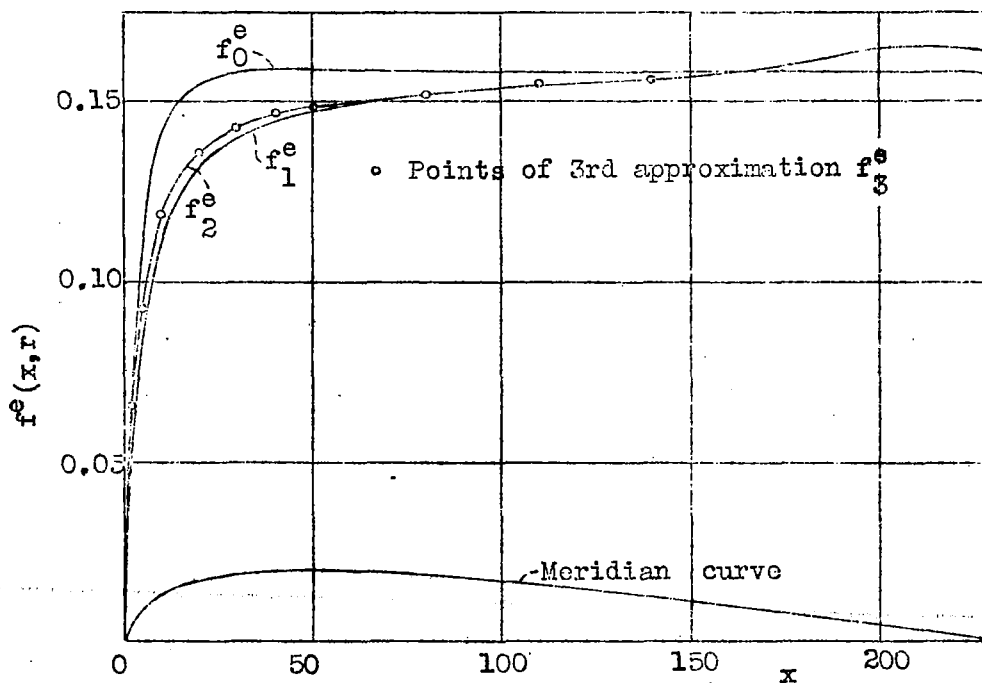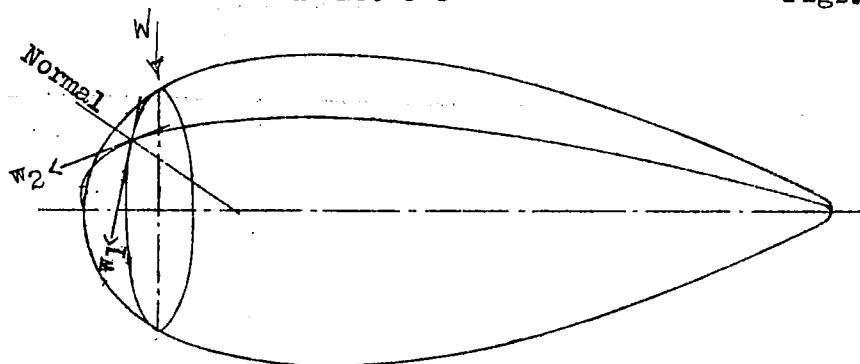

Fig. 20   Surface superposition of Fuhrmann shape No. I

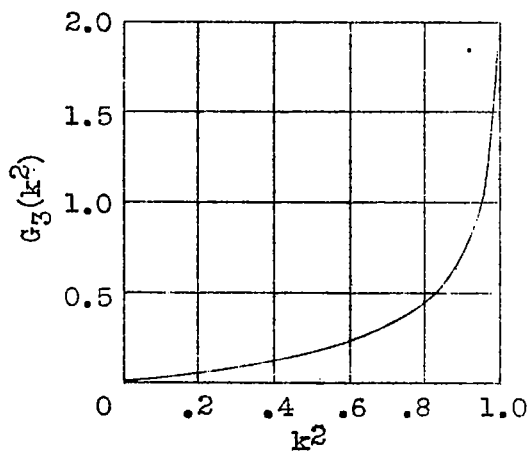Fig. 21 Velocity components on surface of body.
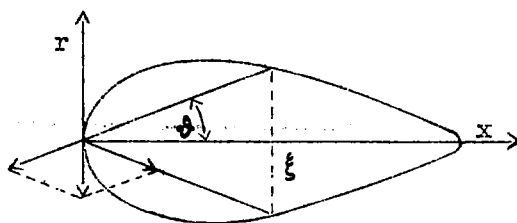

Fig. 22 The $G_3(k^2)$ function

Fig. 23 Meridian cut at angle φ